

1      **On the feasibility of monitoring Carbon Monoxide in the  
2      lower troposphere from a constellation of Northern  
3      Hemisphere geostationary satellites. (PART 1)**

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14     **Abstract**

15     By the end of the current decade, there are plans to deploy several geostationary  
16     Earth orbit (GEO) satellite missions for atmospheric composition over North  
17     America, East Asia and Europe with additional missions proposed. Together, these  
18     present the possibility of a constellation of geostationary platforms to achieve  
19     continuous time-resolved high-density observations over continental domains for  
20     mapping pollutant sources and variability at diurnal and local scales. In this paper,  
21     we use a novel approach to sample a very high global resolution model (GEOS-5 at 7  
22     km horizontal resolution) to produce a dataset of synthetic carbon monoxide  
23     pollution observations representative of those potentially obtainable from a GEO  
24     satellite constellation with predicted measurement sensitivities based on current  
25     remote sensing capabilities. Part 1 of this study focuses on the production of  
26     simulated synthetic measurements for air quality OSSEs (Observing System  
27     Simulation Experiments). We simulate carbon monoxide nadir retrievals using a  
28     technique that provides realistic measurements with very low computational cost.  
29     We discuss the sampling methodology: the projection of footprints and areas of  
30     regard for geostationary geometries over each of the North America, East Asia and  
31     Europe regions; the regression method to simulate measurement sensitivity; and  
32     the measurement error simulation. A detailed analysis of the simulated observation  
33     sensitivity is performed, and limitations of the method are discussed. We also  
34     describe impacts from clouds, showing that the efficiency of an instrument making  
35     atmospheric composition measurements on a geostationary platform is dependent  
36     on the dominant weather regime over a given region and the pixel size resolution.  
37     These results demonstrate the viability of the “instrument simulator” step for an  
38     OSSE to assess the performance of a constellation of geostationary satellites for air  
39     quality measurements. We describe the OSSE results in a follow up paper (Part 2 of  
40     this study).

43

## 44 1. Introduction

45

46 Current satellite observations of tropospheric composition made from low  
 47 Earth orbit (LEO) satellites provide at best one or two measurements each day at  
 48 any given location. Coverage is quasi-global but sparse, often with large  
 49 uncertainties in individual measurements that limit examination of local and  
 50 regional atmospheric composition over short time periods. This has hindered the  
 51 operational uptake of these data for monitoring air quality and population exposure,  
 52 and for initializing and evaluating chemical weather forecasts.

53 By the end of the current decade, there are planned geostationary Earth orbit  
 54 (GEO) satellite missions for atmospheric composition over North America, East Asia  
 55 and Europe, with additional missions proposed (CEOS, 2011). Together, these  
 56 present the possibility of a constellation of GEO platforms to achieve continuous  
 57 time-resolved high-density observations over continental domains for mapping  
 58 pollutant sources and variability. The GEO geometry provides a continuous view of  
 59 the part of the Earth that is below the satellite, enabling measurements many times  
 60 per day that help capture the diurnal evolution of emission sources, tropospheric  
 61 chemistry and pollution transport.

62 The planned GEO missions include the EVI-1/TEMPO (Tropospheric  
 63 Emission: Monitoring of Pollution, Zoogman et al., 2014b) over USA, Sentinel 4/IRS  
 64 over Europe and GEMS over Asia. Each mission has a different primary objective,  
 65 but they share the common goal of monitoring pollutants for atmospheric  
 66 composition and air quality and will have a common measurement capability for  
 67 ozone ( $O_3$ ), nitrogen dioxide ( $NO_2$ ), sulfur dioxide ( $SO_2$ ), formaldehyde ( $HCHO$ ) and  
 68 aerosols, utilizing radiances in the ultraviolet-visible (UV-Vis) spectrum. Planned  
 69 GEO observations of infrared active trace gases of relevance to air quality are  
 70 currently limited to total column carbon monoxide (CO) observations from the  
 71 European IRS instrument, which is originally not driven by atmospheric  
 72 composition applications. However, other IR measurements that could play a part in  
 73 the GEO constellation are being proposed as part of the NASA Decadal Survey GEO-  
 74 CAPE (GEOstationary Coastal and Air Pollution Events) mission, such as the EVI-3  
 75 CHRONOS mission (<https://www2.acd.ucar.edu/chronos>) that would measure CO  
 76 and methane ( $CH_4$ ) using heritage from the Terra/MOPITT (Measurement of  
 77 Pollution in The Troposphere) instrument. Given the effect of nearby emissions and  
 78 transported pollution on local air quality, MOPITT-like CO observations are a good  
 79 candidate for air quality measurements on a GEO platform because the unique  
 80 sensitivity of this platform to pollution in the boundary layer, as well as in the free  
 81 troposphere, allows both vertical and horizontal tracking of pollution transport.

82 Carbon monoxide is a primary pollutant and plays an important role in  
 83 tropospheric chemistry and its sources are both natural and anthropogenic. There  
 84 are two main processes of CO production: incomplete combustion (e.g., industrial  
 85 and urban fossil/bio fuel burning, wildfires and biomass burning); and natural  
 86 chemical production from hydrocarbon oxidation. As an  $O_3$  precursor, CO is also  
 87 important in determining the tropospheric  $O_3$  budget. The principal CO sink is the  
 88 oxidation by hydroxyl (OH) radicals, giving an average CO lifetime of about two  
 89 months dependent on season. With these characteristics, CO serves as a tracer of

90 pollution emissions and transport, and as a proxy for emissions and distributions of  
 91 other species co-emitted with CO but not easily measured. Taken together,  
 92 observations of the full suite of UV-Vis and IR trace gases and aerosols could provide  
 93 the high spatio-temporal resolution continental-scale observations of lower-  
 94 tropospheric pollution needed to monitor, forecast, and manage air quality on a  
 95 daily basis (Edwards et al., 2009; Lahoz et al., 2012; Bowman et al. 2013).

96 Previous GEO observation simulation studies for air quality have assessed  
 97 the potential capabilities of instruments covering the above three continental  
 98 regions separately. Edwards et al. (2009) and Zoogman et al. (2011, 2014ab)  
 99 consider the CONUS (continental US) region and demonstrate the feasibility of using  
 100 observing system simulation experiment (OSSE) studies to help define quantitative  
 101 trace gas measurement requirements in different spectral regions for satellite  
 102 missions and to evaluate the expected performance of proposed observing  
 103 strategies to test the ability of GEO satellite measurements of ozone ( $O_3$ ) and CO.  
 104 Claeysman et al. (2011) and Sellitto et al. (2013b) cover the European region and  
 105 describe the capabilities of a concept nadir thermal infrared sensor proposed for  
 106 deployment onboard a GEO platform to monitor  $O_3$  and CO for air quality purposes  
 107 (MAGEAQ: Monitoring the Atmosphere from Geostationary orbit for European Air  
 108 Quality). Lastly, Zoogman et al. (2014a) assimilate concurrent ozone and CO  
 109 observations and show that geostationary measurement of CO provides significant  
 110 benefit for monitoring ozone.

111 The goal of this study is to evaluate the impact of a future GEO constellation  
 112 on global chemical weather by using the observing system simulation experiment  
 113 (OSSE) technique. Here we primarily consider CO as a good chemical tracer for  
 114 evaluating the impact of a GEO constellation of observations. As described by  
 115 Edwards et al. (2009), chemical OSSEs provide a way of expanding case-specific  
 116 sensitivity studies to assess the impact of future measurements systems. A chemical  
 117 OSSE is composed of several elements (fig. 1 – see also, Timmermans et al., 2014). A  
 118 nature run (NR) (1) represents the atmospheric true state. A complete OSSE needs  
 119 an observation simulator (2) to sample the nature run to produce synthetic  
 120 observations (3). The synthetic observations are then assimilated using a data  
 121 assimilation system (4) into a second atmospheric model, the control run (CR) (5).  
 122 This produces the assimilation run (AR) (6). The impact of concept instrument  
 123 measurements on constraining the modeled state of the atmosphere can then be  
 124 evaluated and assessed (7) by comparing the NR, CR and AR (1, 5 and 6). We  
 125 describe this study in two parts. In the present paper (Part 1) we focus on the NR  
 126 (1), observation simulator (2) and synthetic observations (3). A follow-up article  
 127 (Part 2) will focus on assimilating the simulated measurements and assessing the  
 128 synergies between the different instruments of the constellation by simulating data-  
 129 denial case studies (elements 4 to 7 in fig. 1). This study presents for the first time a  
 130 global GEO constellation OSSE for CO.

131 According to Rodgers (2000), within the remote sensing optimal estimation  
 132 framework one can represent the sensitivity of the retrieved trace gas profile from a  
 133 satellite measurement to the true state of the atmosphere by the averaging kernel  
 134 (AK) function. For accurate observation simulations in an OSSE, we need a full  
 135 radiative transfer model for radiances and their Jacobians (which represent the  
 136 sensitivity of the radiance to the true atmospheric state) to compute the AKs for

137 each atmospheric and surface scene. Since this presents a significant computational  
 138 burden, practical implementations of OSSEs for air quality to date have used  
 139 approximated observation simulators. Some have used specified constant AKs  
 140 (Edwards et al., 2009; Zoogman et al., 2011), or have simplified the AK variability by  
 141 considering only a few scene types (Claeyman et al., 2011). Sellitto et al. (2013a)  
 142 showed that the use of no, or limited scene dependent AK, parameterizations could  
 143 significantly misrepresent the sensitivity of an observing system. Sellitto et al.  
 144 (2013a) also recommend using comprehensive scene-dependent approximations of  
 145 the AKs in cases where the computational cost of a full radiative transfer model is  
 146 too expensive to perform an OSSE study (for example, for a GEO constellation).  
 147 Worden et al. (2013) address this issue by using a multiple regression analysis of  
 148 real satellite observations to estimate scene-dependent averaging kernels, thus  
 149 avoiding the use of a full radiative transfer model. This method allows the fast  
 150 computation of scene-dependent AKs, and the processing of a very large dataset of  
 151 synthetic observations in a short amount of time.

152 Due to the constraint from the NR space and time resolution, approximations  
 153 made to the instrument sampling and horizontal resolution cannot provide  
 154 information at a higher resolution than the nature run (Edwards et al., 2009). One  
 155 should use high space and time resolution NRs to simulate high instrument space  
 156 and time sampling. The planned missions mentioned above would provide less than  
 157 10 km spatial resolution at about every 1 hour. Sellitto et al. (2013b) also  
 158 approximated the observation simulation by not discarding the cloud-contaminated  
 159 measurements, thus leading to a possible overestimation of the GEO instrument  
 160 potential to monitor tropospheric O<sub>3</sub> and pollution features in general. One should  
 161 account for cloud contamination by testing scenarios with variable instrument  
 162 sampling and resolution.

163 In this paper we use the multiple regression analysis of Worden et al. (2013)  
 164 to produce a very large data set representing a GEO constellation of synthetic  
 165 observations for air quality. In section 2, we describe the very high resolution NR  
 166 from the Goddard Earth Observing System Model version 5 (GEOS-5) at 7 km  
 167 horizontal resolution. Section 3 describes the sampling methodology with details on  
 168 the geostationary projection to the surface of the earth, and the multi linear  
 169 regression method with its limitations for predicting averaging kernels and  
 170 estimated observation errors. Section 4 investigates the impacts of clouds on the  
 171 GEO constellation. The effect of horizontal resolution and sampling is discussed.  
 172 Section 5 presents the measurements and a detailed analysis of the simulated  
 173 observation sensitivity (e.g., averaging kernel variability). Section 6 gives a  
 174 summary, conclusions and perspectives.

175  
 176 2. The nature run  
 177

178 The Goddard Earth Observing System Model, Version 5 (GEOS-5, Rienecker et  
 179 al., 2008) is used to provide the NR. The GEOS-5 atmospheric model is a weather-  
 180 and-climate model used for atmospheric analyses, weather forecasts, uncoupled and  
 181 coupled climate simulations and predictions, and for coupled chemistry-climate  
 182 simulations. The NR used for this study covers a 2-year global, non-hydrostatic  
 183 mesoscale simulation for the period 2005-2006. In addition to standard

184 meteorological parameters (wind, temperature, moisture, surface pressure), this  
 185 simulation includes 15 aerosols tracers (dust, sea salt, sulfate, black and organic  
 186 carbon), and O<sub>3</sub>, CO and carbon dioxide (CO<sub>2</sub>) trace gases.

187 The model simulation is driven by prescribed sea-surface temperature and  
 188 sea-ice derived at a horizontal resolution of 0.25 degrees. Biomass burning  
 189 emissions of organic carbon, sulfate, CO and CO<sub>2</sub> are obtained from the Quick Fire  
 190 Emissions Dataset (QFED) version 2.4-r6. The basis of the QFED is the fire radiative  
 191 power (top-down) approach, and it draws on the cloud correction method used in  
 192 the Global Fire Assimilation System (GFAS; Kaiser et al. 2012). Anthropogenic  
 193 emissions of carbon species and aerosols are largely taken from the Emissions  
 194 Database for Global Atmospheric Research (EDGAR; Olivier et al., 1994), which are  
 195 provided annually at a resolution of 0.1 degrees. For CO and CO<sub>2</sub>, EDGAR v4.2  
 196 emissions from 2005 through 2007 were used. For organic and black carbon  
 197 aerosols species, Hemispheric Transport of Air Pollution (HTAP) emissions were  
 198 used.

199 Outputs at 30-minute intervals have been produced at a resolution of 0.0625  
 200 degrees (~7 km) using a cubed-sphere horizontal grid with 72 vertical levels,  
 201 extending from the surface up to 0.01 hPa (~85 km). All details and references  
 202 concerning nature run file specifications, meteorology, chemistry and emissions can  
 203 be found in the NR description documents at:  
 204 <http://gmao.gsfc.nasa.gov/projects/G5NR/>

205 In this study we focus on July 2006. Figure 2 shows the CO total column  
 206 provided by the NR for 15 July 2006 at 15:00 UT. This map shows the ability of the  
 207 NR to represent the high variability of CO fields at a global scale. We display typical  
 208 and expected CO values: very high values (above 4.10<sup>18</sup> molecules/cm<sup>2</sup>) over central  
 209 Africa due to biomass burning; high values (around 3.10<sup>18</sup> molecules/cm<sup>2</sup>) over  
 210 dense populated areas due to anthropogenic emissions. The NR total columns of CO  
 211 also clearly show long-range transport patterns of CO from anthropogenic and  
 212 biomass-burning sources across the oceans of the Northern Hemisphere (NH) and  
 213 Southern Hemisphere (SH), respectively.

214 Figure 3 shows the July 2006 average of surface CO values over the three  
 215 regions of interest (North America - CONUS, Europe and Eastern Asia). The NR  
 216 shows realistic horizontal CO variability due to the very high space and time  
 217 resolutions of the simulations. Emissions from cities from small to large size are  
 218 clearly identifiable. Transport infrastructure such as roads (eastern US in figure 3.a)  
 219 and ship routes (China sea in figure 3.c) are also visible. In this study we use the NR  
 220 model output variables, both the chemical parameters (CO quantities) and the  
 221 meteorological parameters (not shown), to predict averaging kernels for simulated  
 222 observations in the GEO constellation. This is done for each of the CONUS, Europe  
 223 and Eastern Asia regions of interest.

224  
 225 3. Sampling methodology

226  
 227 3.1. Geometry of measurements

228 We constructed three GEO instrument simulators over the three regions of  
 229 interest defined immediately above using the methodology described in Worden et

231 al. (2013). Footprints of the instruments are defined as a GEO projection on the  
 232 globe. We defined  $x$  (along the parallel from the sub-satellite point) and  $y$  (along the  
 233 meridian from the sub-satellite point) at regularly spaced scanning angles (in  
 234 degrees). The GEO projection consists of projecting these angles from the GEO  
 235 platform to the surface of the earth to obtain the corresponding longitudes and  
 236 latitudes of the footprints. We have the following relationship between viewing  
 237 angles at the satellite location and latitude, longitude position on the earth surface:

238

$$239 \quad lon = \tan^{-1}(s_1/s_2) + sub\_lon \quad (1)$$

240

$$241 \quad lat = \tan^{-1}(p_2(s_3/s_{xy})) \quad (2)$$

242

243 where  $sub\_lon$  is the sub-satellite point longitude and:

244

$$245 \quad s_1 = p_1 - s_n \cos x \cos y$$

246

$$246 \quad s_2 = s_n \sin x \cos y$$

247

$$247 \quad s_3 = -s_n \sin y$$

248

$$248 \quad s_{xy} = \sqrt{s_1^2 + s_2^2}$$

249

$$249 \quad s_d = \sqrt{(p_1 \cos x \cos y)^2 - ((\cos y)^2 + p_2(\sin y)^2)p_3}$$

250

251  $p_1 = 42164$  km, the altitude of a GEO platform from the center of the earth  
 252  $p_2 = 1.006803$  is the ratio of the earth radius at the equator and at the pole ( $p_2 =$   
 253  $r_{eq}/r_{po}$ ).  
 254  $p_3 = p_1^2 - r_{eq}^2$

255

256 These equations follow from the methods provided in the technical  
 257 document EUMETSAT (2011) and sketch of figure 4.d should be consulted to  
 258 understand the above formulas. Projecting the regularly spaced instrument viewing  
 259 angles onto the surface of the earth (figure 4.b) results in GEO instrument footprints  
 260 with non-regular latitude and longitude spacing. GEO instruments then have a non-  
 261 uniform horizontal resolution: the footprint density per surface area decreases as  
 262 the measurements go outward from the sub-satellite point (figure 4.c). The GEO-  
 263 CAPE concept mission (Fishman et al., 2012) requires hourly measurements with a  
 264 spatial resolution in the order of 5 to 10 km and a measurement domain of at least  
 265 5000 km. Table 1 gives an overview of the characteristics of the three instruments  
 266 that we call hereafter GEO-US (over CONUS), GEO-EU (over Europe) and GEO-AS  
 267 (over Eastern Asia). We set the scanning angles of the three instruments to have a  
 268 horizontal resolution under 10 km ( $0.1^\circ$ ) in the approximate middle of the

269 measurement domain (i.e., sub-longitude and the mean of latitudes at the sub-  
 270 longitude). Figure 4.a. shows the measurement domains of the GEO constellation.  
 271 Areas of coverage have different shapes due to the latitudinal extent of continents;  
 272 GEO-EU has more of a latitudinal extent compared to GEO-US, which has to cover a  
 273 wider longitude range. GEO-AS has been designed as a compromise solution  
 274 between measurements over Chinese mega-cities and measurements over Korea  
 275 and Japan.

276

## 277 3.2 Carbon monoxide instrument simulator

278

279 In this study, we assume characteristics of the CO measurements of the  
 280 troposphere similar to those of the Terra/MOPITT (Measurement of Pollution in the  
 281 Troposphere) instrument (Drummond, 1992). The last version of the retrieved CO  
 282 product version 5 (Deeter et al., 2013) uses a multispectral approach utilizing near-  
 283 visible infrared (NIR) solar backscatter signals at 2.3 microns and thermal infrared  
 284 (TIR) emission signals from the Earth surface and atmosphere at 4.6 microns. This  
 285 approach provides enhanced measurement sensitivity to near-surface CO  
 286 concentrations and allows the possibility of retrieving CO profile information to  
 287 separate CO in the planetary boundary layer and free troposphere (Worden et al.,  
 288 2010). This is a requirement for the GEO-CAPE concept mission (Fishman et al.,  
 289 2012) and it is generally desirable for air quality space remote observations to  
 290 distinguish between local emissions and transported pollution at a given location  
 291 (Lahoz et al., 2012). In the case of MOPITT, the combination of the TIR and NIR  
 292 radiances significantly improves the sensitivity to the lower tropospheric CO for  
 293 daytime land observations. For nighttime land and day/night ocean observations,  
 294 only the TIR radiances contribute to the retrieval.

295 The MOPITT-retrieved CO volume mixing ratios (VMRs) are on 10 pressure  
 296 levels (surface, 900, 800, 700, 600, 500, 400, 300, 200, 100 hPa). Each retrieved  
 297 level is representative of the layer content defined by the level value itself and the  
 298 level above. The top most level extends from 100 hPa to 50 hPa. The retrieved CO  
 299 profile  $y_r$  can be related to the true atmospheric state  $y_t$  with the following linear  
 300 relationship:

301

302 
$$y_r = y_a + A(y_t - y_a) + \epsilon \quad (3)$$
 303

304 In Eq. (3)  $y_t$  is the true atmospheric CO profile state (in  $\log_{10}(\text{VMR})$ ) and  $y_a$  is the a-  
 305 priori state vector (in  $\log_{10}(\text{VMR})$ ) derived from a monthly mean climatological  
 306 profile from the MOZART-4 (Model for Ozone and Related chemical Tracers, version  
 307 4) chemical transport model (Emmons et al., 2010). The random error  $\epsilon$  (in  
 308  $\log_{10}(\text{VMR})$ ) is simulated using the retrieval noise, and  $A$  is the retrieval AK matrix  
 309 (see section 3.3). (The  $y_r$  retrieved profile obtained is then converted from  
 310  $\log_{10}(\text{VMR})$  to VMR for the final data product).

311 Figure 5 shows two representative MOPITT AKs. The sensitivity of the  
 312 MOPITT instrument to near-surface CO varies according to different surface types  
 313 and atmospheric conditions. The left panel of Fig. 5 shows a typical AK for a daytime  
 314 measurement over land with enhanced sensitivity toward the surface. The right  
 315 panel of Fig. 5 shows a typical AK for a TIR-only ocean or nighttime measurement

316 over land with low sensitivity in the lowermost troposphere. A useful quantity  
 317 indicating the information content of a measurement is the degrees of freedom for  
 318 signal (DFS), given by  $tr(\mathbf{A})$  (Rodgers, 2000). Higher DFS values indicate more  
 319 sensitivity of the retrieval to the true profile.

320 To diagnose the sensitivity of the measurement to the lowest layers, DFS can  
 321 be calculated over the three lowest levels (Surface to 700 hPa) as  $DFS_{0,3} = \sum_{i=1}^3 A_{ii}$ .  
 322 In Figure 5 the DFS ( $DFS_{0,3}$ ) is 1.9 (0.7) and 1.5 (0.2) for land-day and ocean-night  
 323 measurements, respectively. We can see that MOPITT sensitivity toward the surface  
 324 ( $DFS_{0,3}$ ) is scene dependent. That is, it depends on various land and atmospheric  
 325 parameters (i.e., nature of the surface and current state of the atmosphere at a given  
 326 time) that control, among other things, the surface-atmospheric thermal contrast,  
 327 i.e., the difference between the surface temperature and the atmospheric  
 328 temperature profile.

329  
 330 3.3 Simulated retrieval method.  
 331

332 Worden et al. (2013) investigated the CO retrieval error resulting from the  
 333 use of a single average AK in an observation simulator compared using the true  
 334 retrieval AKs. They further developed a scene-dependent AK prediction tool capable  
 335 of approximating the true AK with a significant reduction in retrieved CO error  
 336 compared to using a single average AK. This AK prediction tool allows us to produce  
 337 a large amount of simulated data over months in an efficient manner. One month of  
 338 data for a GEO constellation (i.e., around 200 million profiles) can be produced in  
 339 less than 12 hours.

340 The method of Worden et al. employs a multiple regression approach for  
 341 deriving scene dependent AKs using predictors based on state parameters from the  
 342 NR. The main parameters used are: CO concentration, temperature, specific  
 343 humidity and pressure (see table 2). The method is based on the computation of the  
 344 singular value decomposition (SVD) of the AK matrix. Given an AK matrix  $\mathbf{A}$ , we  
 345 compute the SVD by means of:

346  
 347 
$$\mathbf{A} = \mathbf{U}\Lambda\mathbf{V}^T \quad (4)$$
  
 348

349 where the columns of  $\mathbf{U}$  and  $\mathbf{V}$  are the left and right singular vectors respectively,  
 350 and the elements of  $\Lambda$  (a diagonal matrix) are the singular values. Since the first two  
 351 singular vectors account for 95% of the variability of MOPITT CO AKs on average  
 352 and the first three singular vectors account for 99.995 %, the method retains the  
 353 first three ranked singular vectors. For a complete description of the SVD technique,  
 354 numerical examples and software used please refer to Worden et al., 2013. We then  
 355 calculate the three first singular vectors and values using multiple regression. For  
 356 example,

357  
 358 
$$U_{ij} = c_{ij} + \sum_{k=1}^N a_{ijk} x_{jk} \quad (5)$$
  
 359

360 with dimensions of:  $i$  singular vectors,  $j$  pressure levels, and  $k$  predictors. The  
 361 parameters are:  $c$ , a constant;  $a$ , regressions coefficients; and  $x$ , predictors. We used  
 362 twelve predictors ( $N=12$ ) and have defined eleven different training sets (containing

363 the  $\mathbf{a}$  coefficients) for the geographical regions of interest. Only a single training set  
 364 can be used in the regression calculation. The predictors and training sets are listed  
 365 in Table 2. Worden et al. (2013) selected the predictors based on their importance  
 366 in the regression technique for parameterizing MOPITT forward model  
 367 transmittances of Edwards et al. (1999). The training sets are derived from a multi  
 368 linear fit using real MOPITT observations. The training set period is the entire year  
 369 2006. Once an AK matrix  $\mathbf{A}$  is predicted, the simulated observation profile from the  
 370 NR can be computed using the retrieval equation:

$$372 \quad \mathbf{y}_r = \mathbf{y}_a + \mathbf{A}(\mathbf{y}_{NR} - \mathbf{y}_a) + \boldsymbol{\varepsilon} \quad (6)$$

374 with  $\mathbf{y}_{NR}$ , the NR profile sampled at the MOPITT vertical resolution, replacing the  
 375 true state profile  $\mathbf{y}_t$  in equation 3. Because MOPITT retrieved values express a CO  
 376 quantity over a pressure layer, we compute a weighted average using the pressure  
 377 thickness of the GEOS-5 vertical CO levels mapped onto the MOPITT grid to produce  
 378  $\mathbf{y}_{NR}$ .

### 380 3.4. Training set method limitations

382 In section 3.3 we applied the method described Worden et al. (2013), to  
 383 reconstruct the averaging kernel matrix. In order to utilize the multi linear  
 384 regression (equation 5), we need pre-calculated coefficients ( $\mathbf{a}_{0,N}$ ) from a multi-  
 385 regression fit derived from real MOPITT observations, that we call training sets  
 386 given in Table 2. In some cases, mostly over the CONUS and Asian megacities, very  
 387 high CO profile concentrations and total CO column amount values can extend  
 388 beyond the boundary values of the data set used to build the training set and hence  
 389 beyond the boundary values of the training set itself. Because of the near linear  
 390 relationship between predictors and predicted AKs (equation 4 and 5), using  
 391 predictors from the model with values that are outside the training set distributions  
 392 may lead to unphysical averaging kernel values, e.g., strong negative values or  
 393 values above unity. This is most likely the case for the CO predictors (CO profile and  
 394 CO total column). In order to prevent predictors that are outside the training set  
 395 range and not to discard a significant amount of simulated observation over  
 396 polluted areas we reduce the CO profile predictor as follows. We calculate the mean  
 397 ( $\mu$ ) and standard deviation ( $\sigma$ ) of the CO profile training sets. If the predicted CO  
 398 profile values ( $p$ ) are above  $\mu+2\sigma$ , the new predictor ( $p'$ ) is then calculated as  
 399 follows:

$$400 \quad 401 \quad p' = (1-\gamma)(\mu+2\sigma) + \gamma p \quad (7)$$

402 where  $\gamma$  is a weighting coefficient ranging between 0 and 1. Then the scaled CO  
 403 profile is used to recalculate related CO predictors (CO column,  $\text{Cos}(\theta_{sza})/\log_{10}\text{CO}(z)$   
 404 and  $dT(z)/\log_{10}\text{CO}(z)$ ). This allows the simulator to produce reasonable variability  
 405 in measurement sensitivity while still including the high CO cases and without  
 406 generating unphysical averaging kernels. Sensitivity tests during extreme pollution  
 407 events have shown that using  $\gamma>0.2$  produces an unacceptably high frequency of

409 unrealistic averaging kernel functions. In order to have a robust observation  
 410 simulator which does not produce unphysical averaging kernel values we use  $\gamma=0.1$ .  
 411

412 3.5. Simulated error method  
 413

414 The regression method described above does not account for simulating  
 415 measurement error (represented by the retrieval error covariance matrix) and  
 416 retrieval noise. In order to simulate the error terms, we use the relationships  
 417 between the AK matrix and the associated retrieval errors terms (Rodgers, 2000).  
 418 The associated retrieval noise  $\varepsilon$  is defined using the retrieval noise covariance  
 419 matrix  $\mathbf{C}_n$ , derived from the retrieval error covariance matrix  $\mathbf{C}_x$ . Where  $\varepsilon$  is the  
 420 vector containing the square root of the diagonal elements of  $\mathbf{C}_n$ , and  $y_{err}$  the  
 421 vector containing the square root of the diagonal elements of  $\mathbf{C}_x$ . The retrieval error  
 422 covariance matrix  $\mathbf{C}_x$  can be decomposed as the sum of two matrices (Deeter et al.  
 423 2011):  
 424

- 425 • A smoothing error covariance matrix  $\mathbf{C}_s$  that describes the expected error  
 426 arising from differences between the true profile and retrieved profile, and  
 427 which are due to the characteristics of the weighting functions and the  
 428 influence of the a priori covariance matrix.
- 429 • A retrieval noise error covariance matrix  $\mathbf{C}_n$  that quantifies the expected  
 430 errors due to errors in the radiances.

431 Then

$$434 \quad \mathbf{C}_x = \mathbf{C}_s + \mathbf{C}_n \quad (8)$$

436 with  $\mathbf{C}_s$  approximated using the a priori covariance matrix  $\mathbf{C}_a$ , as follows  
 437

$$438 \quad \mathbf{C}_s = (\mathbf{I} - \mathbf{A})\mathbf{C}_a(\mathbf{I} - \mathbf{A})^T \quad (9)$$

440 and  $\mathbf{C}_x$  directly calculable from  $\mathbf{C}_a$  and  $\mathbf{A}$   
 441

$$442 \quad \mathbf{C}_x = (\mathbf{I} - \mathbf{A})\mathbf{C}_a \quad (10)$$

444 so that

$$446 \quad \mathbf{C}_n = \mathbf{C}_x(\mathbf{I} - (\mathbf{I} - \mathbf{A})^T) \quad (11)$$

449  $\mathbf{C}_n$  is mostly lower than  $\mathbf{C}_s$  but not negligible (see section 5.2 and figure 11).  
 450 Relatively to  $\mathbf{C}_x$ ,  $\mathbf{C}_n$  will increase if  $\mathbf{C}_s$  decrease (if  $\mathbf{A}$  tends to be the identity  $\mathbf{I}$ ). We  
 451 define  $\mathbf{C}_a$  as for the MOPITT v4 and v5 products (Deeter et al., 2010). The a priori  
 452 covariance matrix  $\mathbf{C}_a$  incorporates the same variance value  $C_0$  at all levels, with a  
 453 constant correlation height  $P_c$  over a pressure level  $p$  defining the off-diagonal  
 454 elements. Thus,  
 455

456  $C_{a,ij} = C_0 e^{-(p_i - p_j)/P_c}^2$  (12)

457

458 with  $P_c=100$  hPa and  $C_0=(0.3 \log_{10}e)^2$ . In order to simulate the random error  $\varepsilon$ , we  
459 add a pseudo-random noise on each nature run sampled by a predicted AK:

460

461  $\varepsilon = \mathbf{y}_r \mathbf{C}_n^{1/2} \circ \mathcal{N}(0, \mathbf{I})$  (13)

462

463 where  $\circ$  denotes the Schur product and  $\mathcal{N}(0, \mathbf{I})$  a matrix following a normal  
464 distribution of means equal 0 and standard deviation equal the identity matrix  $\mathbf{I}$ . We  
465 also calculate the retrieval error profile as follows:

466

467  $y_{err,i} = y_{r,i} \mathbf{C}_{x,i,i}^{1/2}$  (14)

468

469 Because the smoothing error  $\mathbf{C}_s$  mostly dominates on the error budget (equation 8),  
470 the impact of the random error  $\varepsilon$  is low compared to the retrieval error profile and  
471 hence the accuracy of the retrievals are not significantly impacted.

472

#### 473 4. Impact of clouds

474

475 Under cloudy conditions, the simplest approach for MOPITT-like  
476 measurements on a GEO platform would be simply to discard cloudy pixels and not  
477 perform retrievals. It is thus important to assess the impact of the cloud coverage on  
478 GEO measurements. In this study, a scene is considered clear when the interpolated  
479 cloud fraction from the NR is lower than 5% of a single footprint. This is the clear-  
480 sky condition used operationally with real MOPITT measurements. Cloud  
481 contaminated footprints with greater than 5% of cloud fraction would be discarded.  
482 Clouds properties are not used to predict AK variability. Figure 6 presents the ratio  
483 of cloud free pixels, over the month of July 2006 for the constellation. The ratio of  
484 cloud free pixels is the number of cloud free observations divided by the total  
485 number of possible observations (i.e., one per hour during one month) for a given  
486 pixel. Figure 9 gives an idea of instantaneous instrument coverage with a 5% cloud  
487 fraction threshold. The GEO-EU displays very few cloud-contaminated areas  
488 whereas the GEO-AS has very few cloud free areas.

489

490 The cloud-free ratio geographical distribution shows differences between  
491 intra- and inter-continental regions. On average, GEO-EU has the highest ratio  
492 (60%) followed by GEO-US (40%) and GEO-AS (20%). Strong variations of the ratio  
493 are also observed for different weather regimes within each measurement domain.  
494 Mediterranean weather regimes such as western CONUS and the entire  
495 Mediterranean basin exhibit higher ratios, above 80%. Conversely, oceanic,  
496 subtropical and tropical regimes such as northern Europe, southern CONUS and  
497 southeastern Asia have lower ratios, below 20%. Over the GEO-AS field of view,  
498 Korea and Japan exhibit very low ratios around 10% due to East Asian monsoon  
499 effects that provide persistent convective cloud coverage.

500

501 The value of the cloud free ratio depends on the spatial resolution of the  
observation (pixel size) and the cloud fraction threshold used. Figure 7 displays  
results of sensitivity tests on pixel size and cloud fraction threshold. We assume that

502 the lowest pixel size simulated is 7 km due to the model horizontal resolution. We  
 503 can then increase the pixel size by averaging contiguous grid cells. It is shown here  
 504 that with a given cloud fraction threshold, increasing the pixel size reduces the  
 505 average cloud free ratio. We perform tests for varying cloud fraction thresholds to  
 506 calibrate the assimilated data product. Variations of the cloud free ratio following  
 507 variations in the cloud fraction threshold and the pixel size show the same patterns  
 508 (but with a different range of values) for the three instruments of the constellation.

509 To explain these patterns we display a specific case (figure 8) as an example  
 510 of how the observed coverage changes with the two varying parameters. The case  
 511 study presented shows two typical horizontal cloud structures: one of high  
 512 granularity located over the eastern part of the plot, which is identified as  
 513 convective structures, and the other of low granularity located on the northwest  
 514 part of the plot, which is identified as a cold air front. Over low granularity areas,  
 515 decreasing the cloud fraction threshold will not increase the cloud-contaminated  
 516 area as much as it does over the high granularity areas. As an idealized example, one  
 517 can imagine adding pixels around four single separated sparse pixels (a granular  
 518 structure) and adding a pixel around a four-by-four pixel area (a non-granular  
 519 structure). In the first case, there will be 8 pixels around each of the 4 original  
 520 pixels, making a total of 32 additional pixels. In the second case 12 additional pixels  
 521 will surround a 2 by 2 square. The increase in area will be larger with the granular  
 522 structure than with the non-granular structure.

523 In the more realistic case of our observation simulations, granularity can  
 524 vary at different scales and at different times. We found that adjusting the cloud  
 525 fraction threshold to 20% for a 42 km pixel size gives comparable statistics of cloud  
 526 coverage as with the 5% threshold for a 7 km pixel size (see section 5.3 and figure  
 527 12).

## 528 529 5. Simulated GEO constellation measurements

### 530 531 5.1. Simulated sensitivity analysis

532  
 533 Figure 9 displays the maps of sampled Surface-700 hPa NR and retrieved  
 534 partial columns and associated  $DFS_{0,3}$  for the GEO constellation. Looking at  $DFS_{0,3}$   
 535 maps first shows that the observation simulator reproduces the variability of  
 536 measurement sensitivity over the satellite measurement domains. The maps are  
 537 snapshots during daytime, and show strong differences in  $DFS_{0,3}$  between sea and  
 538 land due to the different AK training sets used. The land training set simulates multi-  
 539 spectral (TIR+NIR) retrieval AKs in contrast to the sea training set that simulates  
 540 TIR-only retrievals. The  $DFS_{0,3}$  values between land and sea surfaces are in  
 541 agreement with figure 5: instrument sensitivities over land are generally higher  
 542 than over sea, because of the availability of multi-spectral simulated retrievals.  $DFS$   
 543 variability over land, or over sea only, is also simulated using the multi-regression fit  
 544 as described in section 3.3. To describe this variability, we will focus on the analysis  
 545 over land. The most obvious variations of  $DFS_{0,3}$  follow orography. The main reason  
 546 is the reduction in the number of retrieved levels if surface pressure is lower than,  
 547 e.g., 900 hPa. For a constant number of retrieved levels, the variation of the surface  
 548 level layer thickness also plays a significant role (represented by the  $dP$  predictor;

549 see table 2), and a thinner surface layer will contribute less retrieval sensitivity.  
 550 Variations of  $DFS_{0,3}$  can also be correlated to the CO amount in the NR. This  
 551 variability is represented with the CO total column and CO profile predictors. CO  
 552 abundance is a strong predictor of sensitivity due to the use of  $\log_{10}(\text{VMR})$  retrievals  
 553 in MOPITT with corresponding weighting functions that have increasing magnitude  
 554 for increasing VMR (Worden et al., 2013). Finally, the temperature profile and  
 555 thermal contrast ( $dT$ ) play a significant role in the  $DFS_{0,3}$  variability, as expected for  
 556 the TIR contribution in a multispectral instrument. While  $DFS_{0,3}$  depends more on  
 557 predictors such as CO column and  $dP$ , all of the predictors in Table 2 add  
 558 information to the regression fit, as tested in Worden et al. (2013).

559 Figure 10 shows scatter plots of  $DFS_{0,3}$  versus the main DFS variability  
 560 drivers, i.e., parameters mentioned above such as CO concentration,  $dP$  and  $dT$ .  
 561 Night and day values are displayed (blue and red, respectively) showing the  
 562 expected increase of sensitivity during day (simulating a multispectral retrieval)  
 563 compared to night (simulating a TIR-only retrieval). For each region, using an  
 564 alternation of day training sets and night training sets, designed to produce  
 565 multispectral and TIR-only retrieval AKs, respectively, then simulates a diurnal cycle  
 566 of sensitivity. Correlation of  $DFS_{0,3}$  with predictors gives an indication of which  
 567 variables in the NR true state will drive measurement sensitivity. However, this is  
 568 not a deterministic result since actual sensitivity depends on all the predictors,  
 569 together with the distributions of those  
 570 variables as compared to the training set distributions, indicated by the lines in  
 571 Figure 10. Variations in the dependence on predictors can be seen by the different  
 572 distributions in Figure 10 for CONUS, Europe and Asia. Over Asia and Europe,  
 573 overall CO concentrations from NR show significantly lower as compared to the  
 574 training set mean. For Asia, scatter plots do not show any clear dependence between  
 575  $DFS_{0,3}$  and CO concentrations. For Europe, the dependence is more marked during  
 576 daytime. Lower CO predictor values compared to training set mean might lead to  
 577 underestimation of  $DFS_{0,3}$ , however it fits a realistic range of values (from 0.25 to  
 578 0.5).

579  
 580 5.2 NR sampling and error budget  
 581

582 The difference between the NR CO and the retrieved CO shows higher NR  
 583 values than in the simulated retrievals (fig. 9). Retrieved values can be close to the  
 584 NR if sensitivity (DFS) is high enough and/or the a priori CO profile is close enough  
 585 to the NR. Cases with strong CO plumes in the NR can be identified in figure 9 over  
 586 Asia (around 35°N and 115°E) and over Europe (around 5°E and 55°N). In the Asian  
 587 case, the plume is very well detected in the synthetic retrieval, because over land  
 588 GEO-AS has a  $DFS_{0,3}$  above 0.5 and a priori profile concentrations close to the NR  
 589 profile (not shown). In the European case, plumes are barely detected because over  
 590 sea the GEO-EU has  $DFS_{0,3}$  below 0.3 and the a priori profile concentrations are far  
 591 from NR values. In general, retrieved CO concentrations are lower than the NR CO  
 592 concentration because a priori values are lower than NR values. In certain cases (see  
 593 fig. 9 for Asia around 110°E and 35°N), the opposite is observed; a priori  
 594 concentrations are higher than the NR. The a priori profile, sampled from a lower  
 595 resolution MOZART-4 climatology (see section 3.2) does not capture the specific NR

596 high-resolution features. Conversely, polluted areas are represented as relatively  
 597 high CO over broad area, which can produce cases where  $y_a$  is higher than  $y_t$ .

598 Figure 11 left panels show scatter plots of NR CO partial columns ( $X_t$ ) versus  
 599 retrieved CO partial columns ( $X_r$ ) with night cases (blue) and day cases (red) over  
 600 land. In general, night  $X_r$  values are farther from the  $X_t$  compared to the day  $X_r$   
 601 values. As explained in section 5.1 and in figure 10,  $DFS_{0,3}$  values are lower during  
 602 night than during day. Lower  $DFS$  will produce  $X_t$  values that are closer to the a  
 603 priori ( $X_a$ ). If  $X_a$  is far from  $X_t$ , the smoothing error ( $X_s$ ) will increase with lower  
 604  $DFS_{0,3}$ . Even if  $DFS_{0,3}$  is high (around 0.7),  $X_s$  can be high if  $X_t$  is very far from  $X_a$ . In  
 605 the case of GEO-US, values spread by 10-20 DU (Dobson Units) around the  $X_t=X_r$   
 606 axis, showing that  $X_a$  can be higher or lower than  $X_t$ . In the case of GEO-EU, the  
 607 spread is lower because  $X_t$  is in general close to  $X_a$ . In the case of GEO-AS,  $X_r$  values  
 608 are mostly lower than  $X_t$  values, showing that  $X_a$  is generally lower than  $X_t$ .

609 Figure 11 right panel displays scatter plots of  $X_s$  (in % relative to  $X_r$ ) versus  
 610 the surface-700hPa partial column retrieval error ( $X_e$ ). We see that  $X_e$  values are in  
 611 the range expected from real MOPITT observations: between 15% and 30%.  
 612 Following equation 8 and 14, diagonal values of  $\mathbf{C}_s$  should be lower or equal to  
 613 diagonal values of  $\mathbf{C}_x$  and hence  $X_s$  should be lower or equal to  $X_e$  (if  $X_s$  is calculated  
 614 as  $X_e$ ). The condition is respected in most of the cases, but some  $X_s$  values are higher  
 615 than  $X_e$ . Again, this happens when  $X_a$  is very distant from  $X_t$ , and due to the fact that  
 616  $X_a$  and  $X_t$  (i.e.,  $y_a$  and  $y_t$ ) are not used in the calculation of the a priori covariance  
 617 matrix (see section 3.5 and equation 9 and 12). The perfect estimate of  $\mathbf{C}_s$  would  
 618 then be:

$$620 \quad \mathbf{C}_s = (\mathbf{I} - \mathbf{A})(\mathbf{y}_a - \mathbf{y}_t)(\mathbf{y}_a - \mathbf{y}_t)^T(\mathbf{I} - \mathbf{A})^T \quad (15)$$

621 This can be estimated for this study since we are assuming the NR is the true state.  
 622 However, for real observations it is not possible to estimate the actual smoothing  
 623 covariance error matrix. Therefore, use of the method described in section 3.5 is  
 624 more realistic, and will provide reasonable error estimates in most cases since  $X_s$   
 625 has generally lower values than  $X_e$ .

### 628 5.3. Reduced resolution simulated observations

630 In part II of this study, we will assimilate the simulated GEO-constellation  
 631 into a global model. We will use the global chemistry – climate model CAM-Chem,  
 632 including its full chemical scheme (Lamarque et al., 2012). State-of-the-art global  
 633 atmospheric chemical models do not have high horizontal resolution. In this second  
 634 part of the study, we use a  $0.9^\circ$  by  $1.25^\circ$  resolution model configuration. Since the  
 635 resolutions of the NR and the simulated observations are much finer than the CAM-  
 636 Chem resolution, we will use the reduced resolution NR ( $0.5^\circ$ , i.e., 42 km  
 637 approximately). The reduced NR simulations are the same as the native NR  
 638 simulations, but the horizontal resolution has been reduced a posteriori (see Da  
 639 Silva et al., 2014).

640 Figure 12 displays the reduced resolution (42 km) simulated observations.  
 641 As explained in section 3.5, because the model resolution is 42 km we assume that  
 642 the pixel size has the same size. To generate an appropriate sampling according to

643 the pixel resolution, we divide by a factor of 5 the number of latitude and longitude  
 644 pixels provided in the table. The left panels show the average surface-700hPa  
 645 retrieved CO column for July 2006. The right panels show the cloud free ratio for  
 646 July 2006. For the cloud fraction threshold, we use 20% to keep the same cloud free  
 647 ratio as for the high-resolution observation simulation, as explained in section 3.5.  
 648 Cloud free ratio maps (figure 12) at low resolution are then very similar to the same  
 649 maps at high resolution (figure 6).

650

651

## 652 6. Conclusion

653

654 This paper is Part 1 of a two-part study. Here, we demonstrate the feasibility  
 655 of simulating a GEO constellation for air quality monitoring, with a focus on CO.  
 656 Three potential instruments are simulated covering the three most populated and  
 657 polluted areas of the world: Continental US (CONUS), Western Europe and Easter  
 658 Asia. We use very high-resolution output ( $0.06^\circ$ , i.e.,  $\sim 7$  km horizontal resolution)  
 659 from the GEOS-5 model as a NR to simulate a MOPITT-like instrument. Instead of  
 660 using a full radiative transfer model to simulate the instrument vertical retrieval  
 661 sensitivity as defined by the AK, we use a novel method described by Worden et al.,  
 662 (2013). This method employs multi-linear regression using predictors (from the  
 663 NR) and training set coefficients (from real MOPITT data) to produce scene-  
 664 dependent AKs, thus allowing a very fast computation of the instrument synthetic  
 665 measurement dataset. Thus, we avoid the computational burden of using a full  
 666 radiative transfer model, allowing the generation of one month of GEO constellation  
 667 data in less than 12 hours. This makes simulation of the GEO constellation  
 668 measurement computationally feasible. The main conclusions of this work are as  
 669 follows:

670

671

1. Instead of using the model resolution as the instrument pixel resolution, and the defined field of view as a simple latitude/longitude rectangle, we present a method to simulate the data using a GEO projection. This gives accurate GEO instrument spatial resolutions and fields of view that vary with latitude and longitude.
2. This paper extends application of the Worden et al., (2013) averaging kernel (AK) prediction method. Realistic variations of potential GEO instrument vertical retrieval sensitivities are simulated. Instrument sensitivities depend on predictors and the main drivers are: surface pressure, CO profile and temperature profile. Rather than using an average AK for fast computation, the observation simulator presented here is able to provide fast computation of AK variability (and its associated retrieval error covariance matrix) at the same time.
3. We discuss limitations of the method used for this study. The very high CO concentrations occurring in the NR over very polluted areas often overreach the training set statistical coverage. In this situation, we use a tuning method to reduce the amplitude of CO variations in the NR.
4. To make the observation simulations as realistic as possible, we account for the impact of clouds. Cloud contamination in the observations is strongly

690 dependent on the instrument spatial resolution and the geographical region  
 691 of interest. The Mediterranean weather regimes show the lowest cloud  
 692 occurrences, whereas subtropical weather might provide comparatively  
 693 lower temporal and spatial sampling for air quality GEO measurements.

694 5. We present case studies for the three measurement domains and show that  
 695 the observation simulation method employed here provides realistic AK  
 696 variability. The degrees of freedom for signal for the lowermost troposphere  
 697 ( $DFS_{0-3}$ ) ranges from 0.2 to 0.7 with significantly larger values over land and  
 698 for day that reflect the enhanced vertical sensitivity possible with  
 699 multispectral retrievals. We simulate small local  $DFS_{0-3}$  variations according  
 700 to surface and atmospheric parameters (e.g., surface pressure, CO profile and  
 701 temperature profile).

702 6. Simulated retrieval errors that are derived from the AK simulation are  
 703 compared to the true smoothing error. Comparisons show that the retrieved  
 704 errors are realistic, being lower than or in the range of the smoothing error.

705  
 706 The next step in this study (Part 2) will be to assimilate the synthetic  
 707 measurement data into a global model. To do so, we present here an additional set of  
 708 simulated observations at a reduced spatial resolution (42 km). This allows an OSSE  
 709 for the potential future prediction system of global air quality with the same  
 710 capabilities for each region of interest: the same models (NR and CR), the same data  
 711 assimilation system (AS) and the same instrument design (observation simulator).  
 712 The goals of Part 2 will be to: (1) assess the ability of the GEO constellation to observe  
 713 the impact of pollutant emissions over each region; (2) look at the importance of  
 714 long-range transport between regions; and (3) investigate the value of the  
 715 measurements from each mission in the GEO constellation, taken individually and  
 716 together.

717  
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Tables:

	GEO-AM	GEO-EU	GEO-AS
Sub_lon	-97°	8.4°	120°
Number x pixels	500	400	400
Number y pixels	230	250	200
X <sub>max</sub>	3.5°	2.4°	3.3°
X <sub>min</sub>	-3.5°	-2.4°	-3.3°
Y <sub>max</sub>	7.2°	8.2°	6.7°
Y <sub>min</sub>	4.2°	5.7°	3.5°

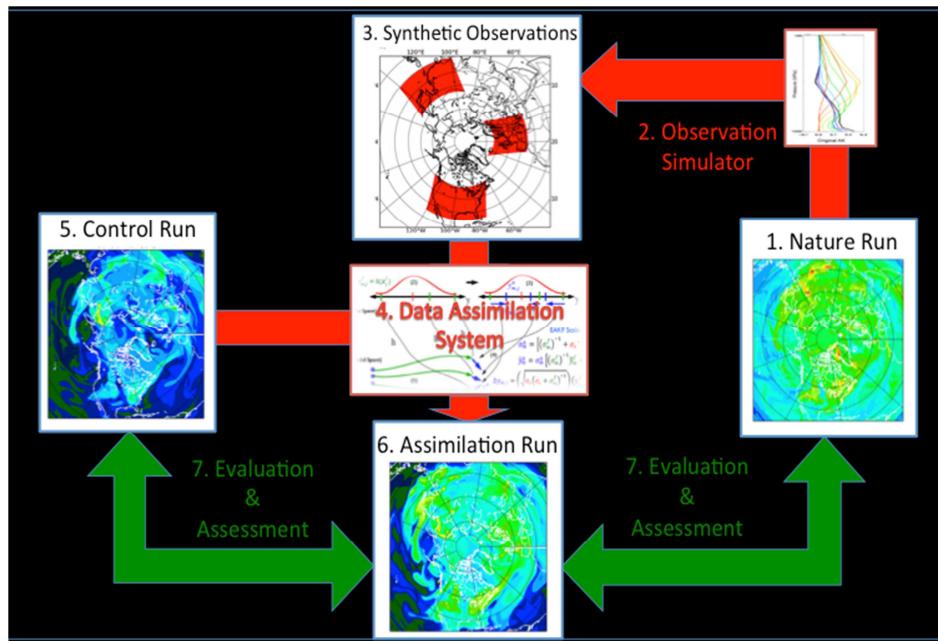
878 **Table 1. GEO-constellation instrument specifications: satellite position,  
879 number of pixels and angles of views.**  
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Predictors	Training sets
$\theta_{sza}$	North Hemisphere Ocean (TIR)
Emissivity	CONUS Day (Psrf>900hPa, TIR+NIR)
Latitude	CONUS Night (Psrf>900hPa, TIR)
Surface temperature	Europe Day (Psrf>900hPa, TIR+NIR)
$dP = P_{surface} - P_{ref}^*$	Europe Night (Psrf>900hPa, TIR)
CO column	Eastern Asia Day (Psrf>900hPa, TIR+NIR)
Water Vapor Q(z)	Eastern Asia Night (Psrf>900hPa, TIR)
CO(z)	N.H. Mountains Day (900hPa>Psrf>800hPa, TIR+NIR)
Thermal contrast $dT(z) = (Tsrf - T(z))$	N. H. Mountains Day (800hPa>Psrf>700hPa, TIR)
$dT(z)^2$	N.H. Mountains Night (900hPa>Psrf>800hPa, TIR+NIR)
$\text{Cos}(\theta_{sza})/\log_{10}CO(z)$	N. H. Mountains Night (800hPa>Psrf>700hPa, TIR)
$dT(z)/\log_{10}CO(z)$	

887 \* $P_{ref}=1000\text{hPa}$ 888  
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899889 **Table 2. Left: List of predictors, right: List of the different training sets used to  
890 produce the geostationary constellation CO measurements. TIR and NIR state  
891 if the training set simulates multispectral or TIR-only retrievals (see text for  
892 details).**  
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## Figures:



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**Figure 1. The chemical OSSE framework. See text for details.**

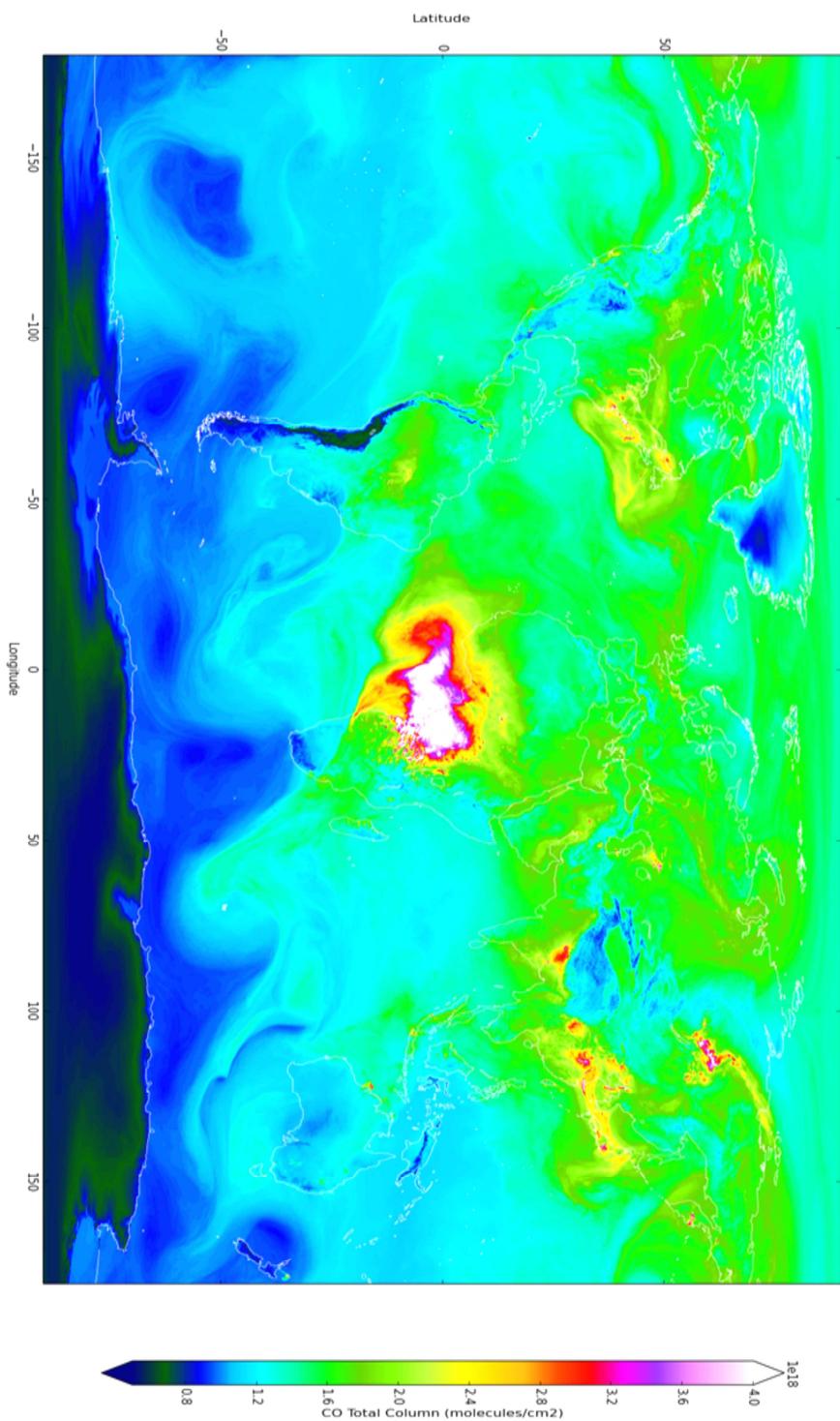
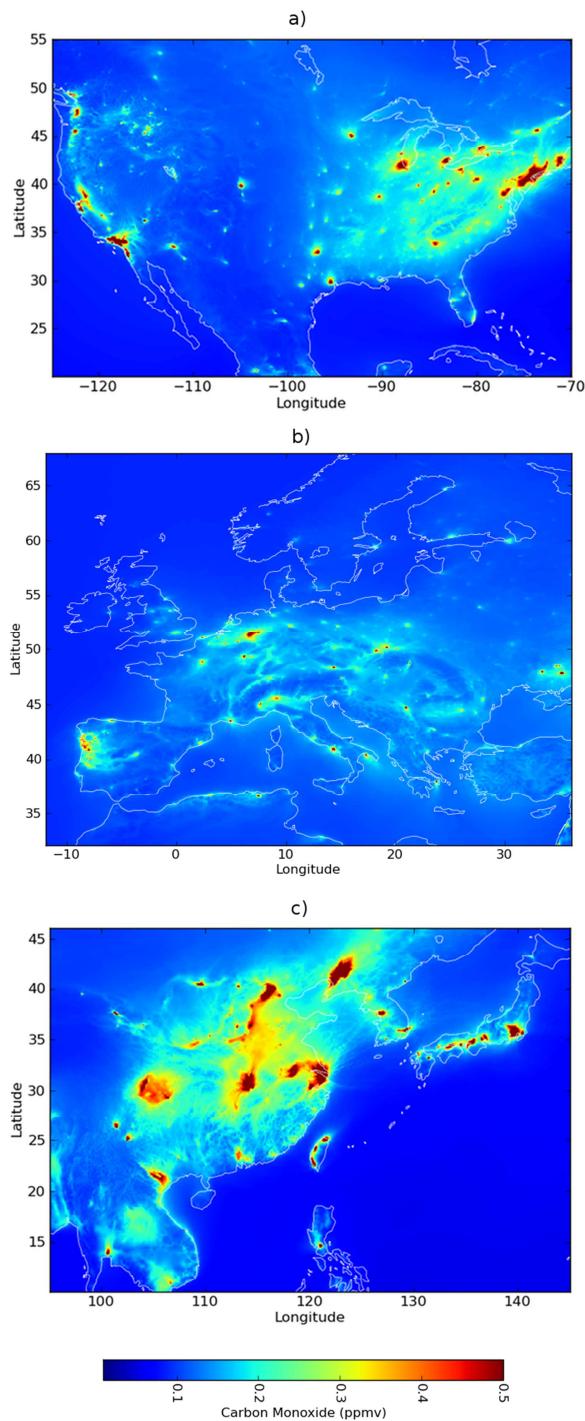
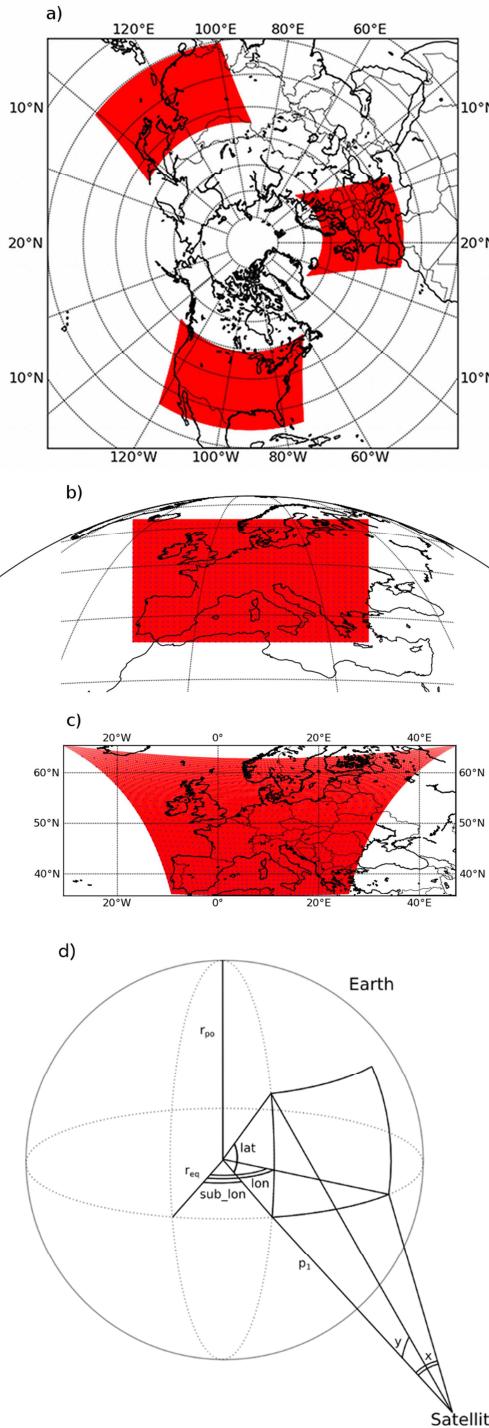


Figure 2. Total carbon monoxide column in molecules/cm<sup>2</sup> from the GEOS-5 7km resolution Nature Run, 15 July 2006 15:00 UT.

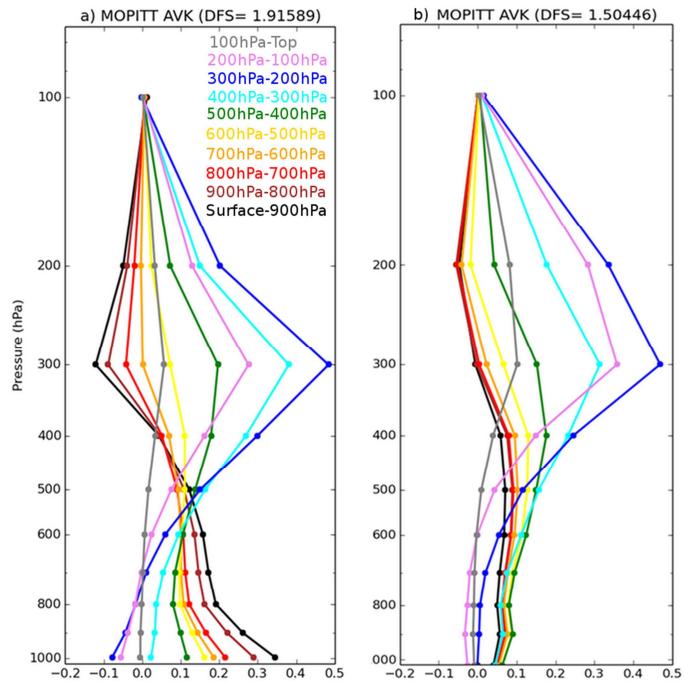


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917 **Figure 3. Surface CO time average during July 2006 over (a) North America, (b)**  
918 **Europe, and (c) Asia.**

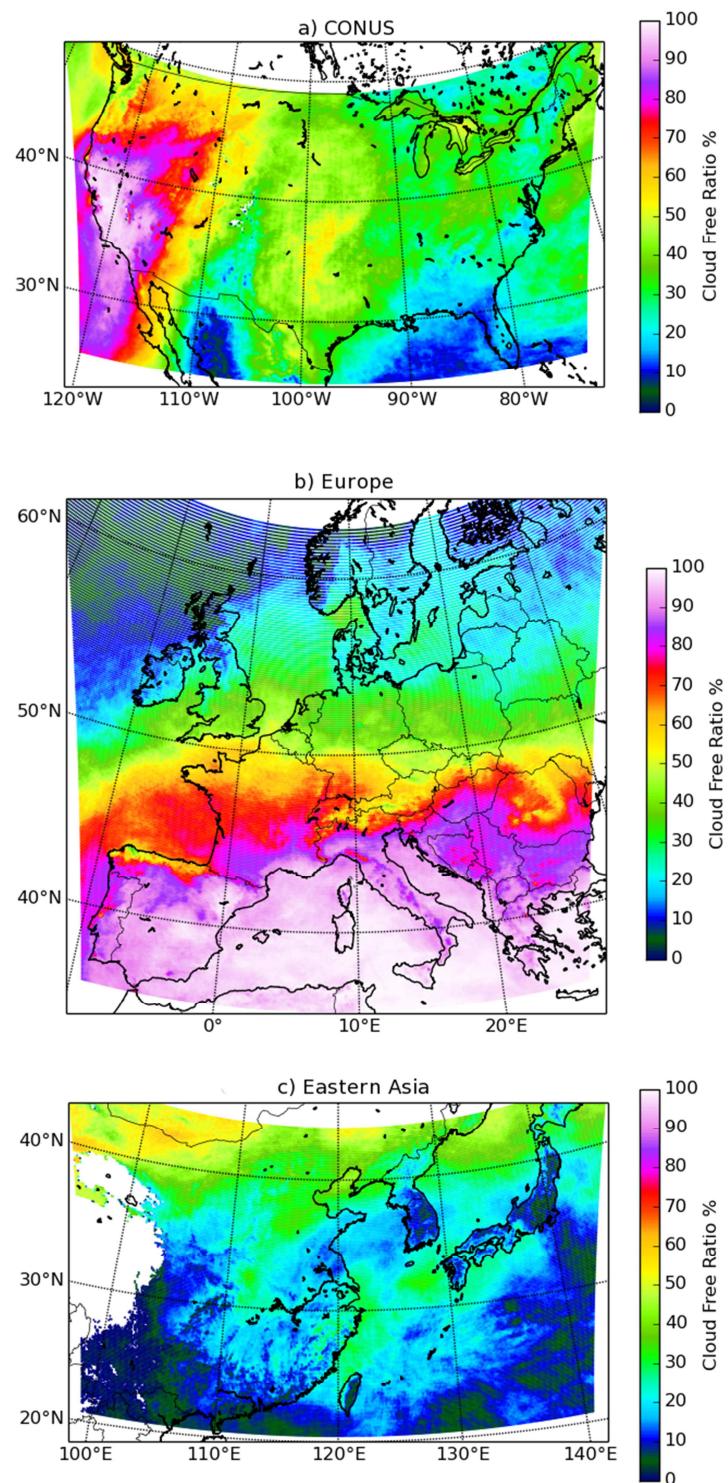
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 922 **Figure 4. a) Geostationary constellation measurement domain a) Polar**  
 923 **projection. b) GEO-EU domain in a geostationary projection, red dots are the**  
 924 **full resolution footprints, purple dots are plotted every 100<sup>th</sup> pixels. c) is the**  
 925 **same as b) but in an equidistant latitude-longitude cylindrical projection. d)**  
 926 **Geometrical sketch of the geostationary projection.**  
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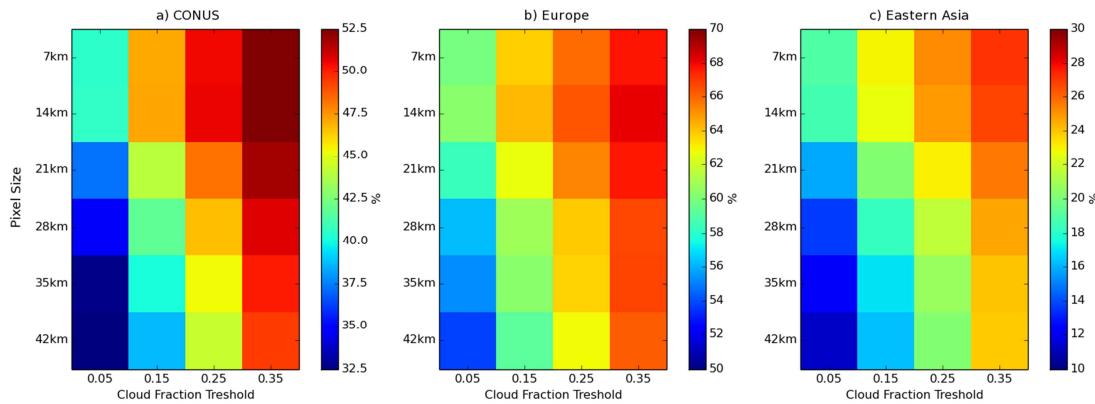


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 929 **Figure 5. Examples of original typical MOPITT averaging kernels (AKs). Left**  
 930 **panel: multispectral day/land AK. Right panel: night/land or sea AK.**

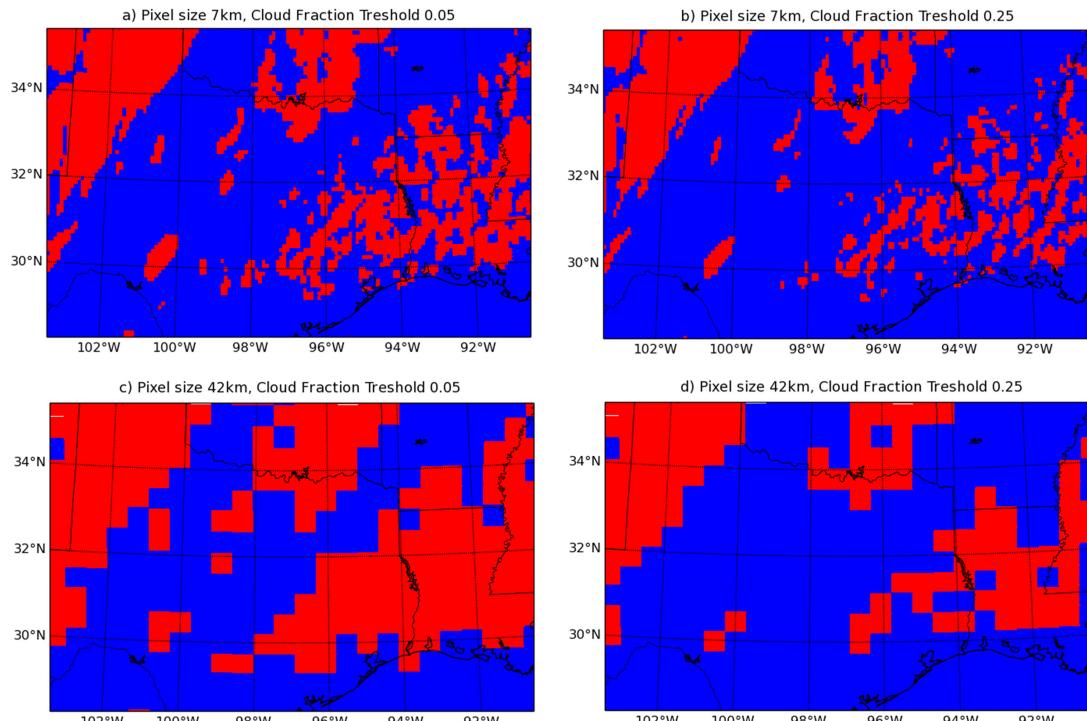


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**Figure 6. Cloud free ratio (%) for the three measurement domains during July 2006.**

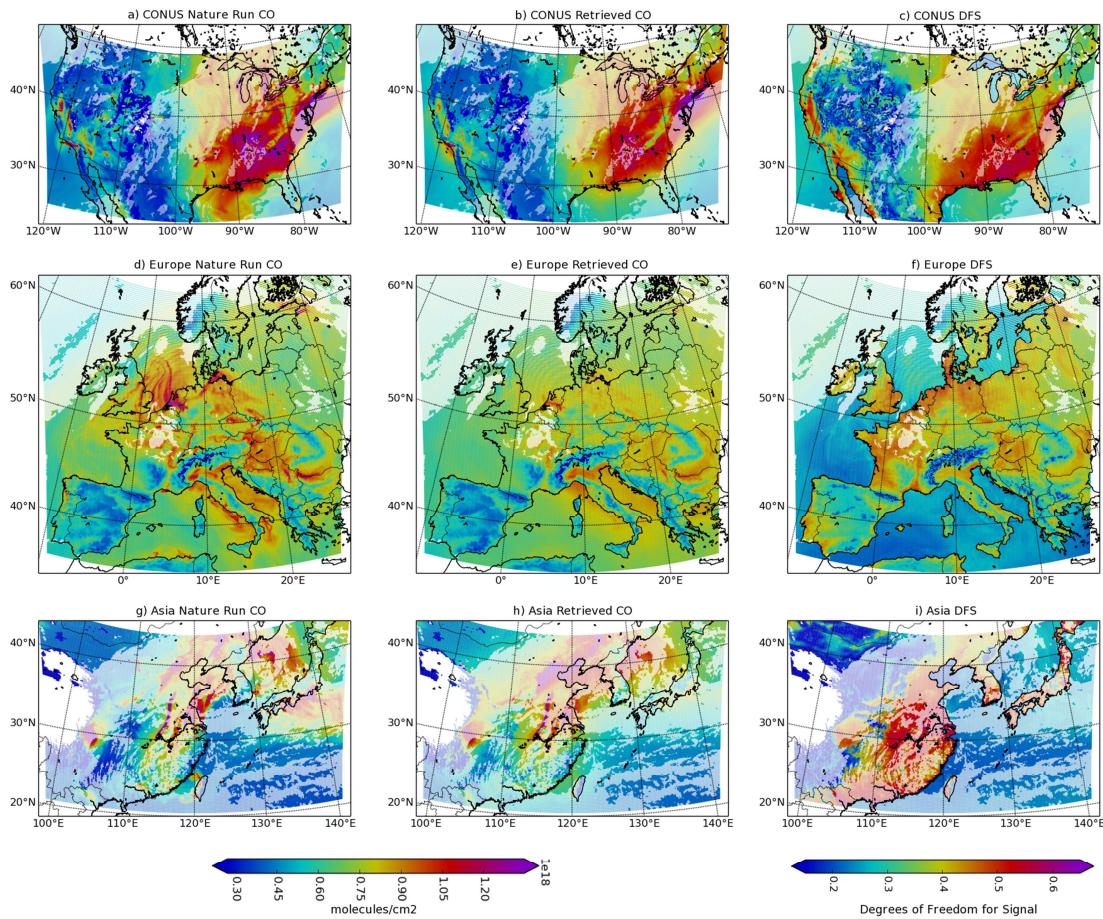


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935 **Figure 7. Sensitivity matrices of the average cloud free ratio (in %) for pixel**  
936 **size versus cloud fraction threshold.**

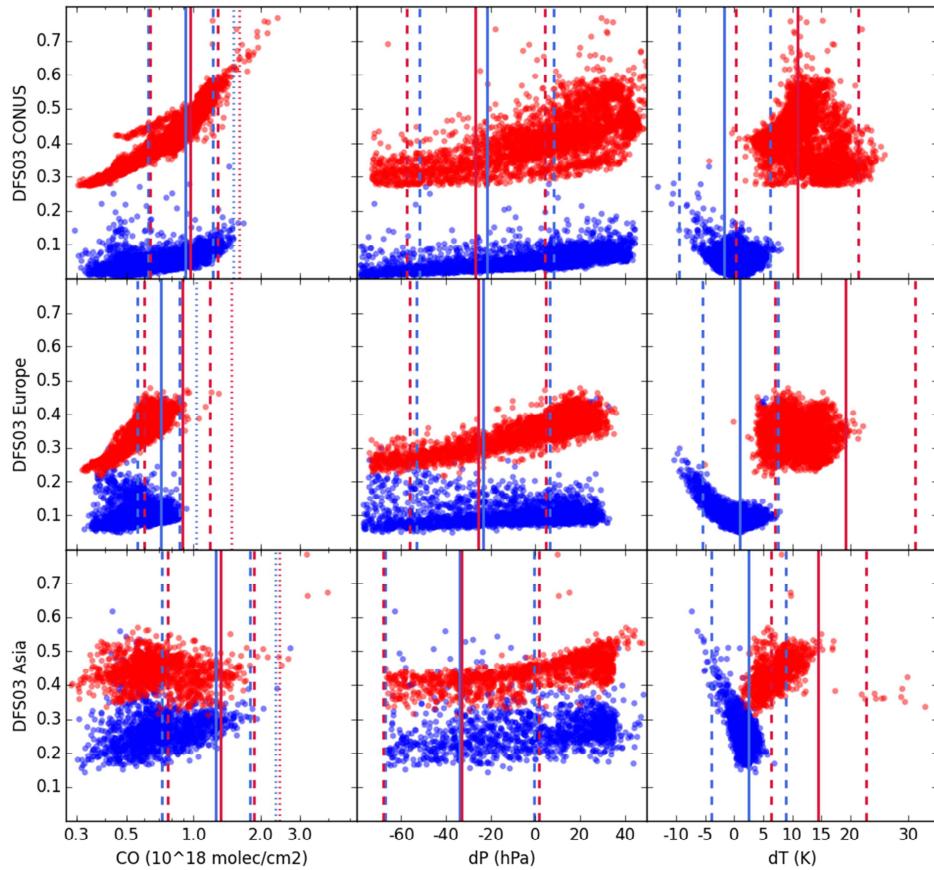


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941 **Figure 8. Examples of cloud detection and ratio of observed area for two**  
942 **different cloud fraction thresholds and two different pixel sizes. Red are cloud**  
943 **contaminated pixels and blue are cloud free pixels. Performed over South East**  
944 **CONUS 5 July 2006 00UT.**

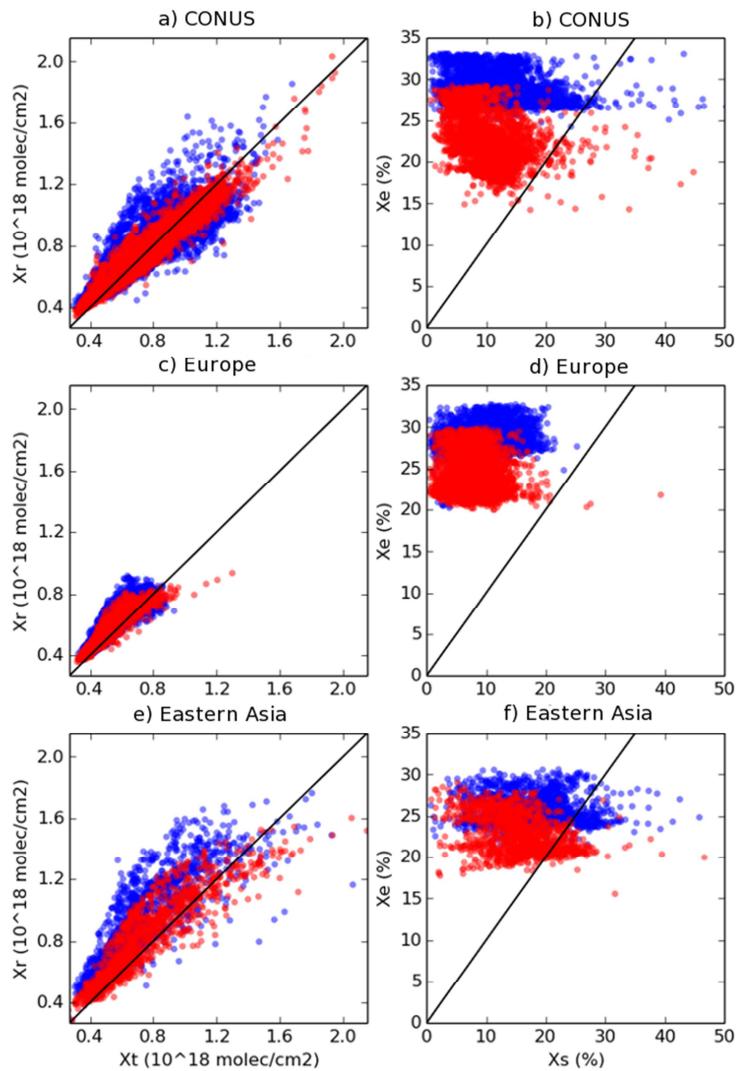
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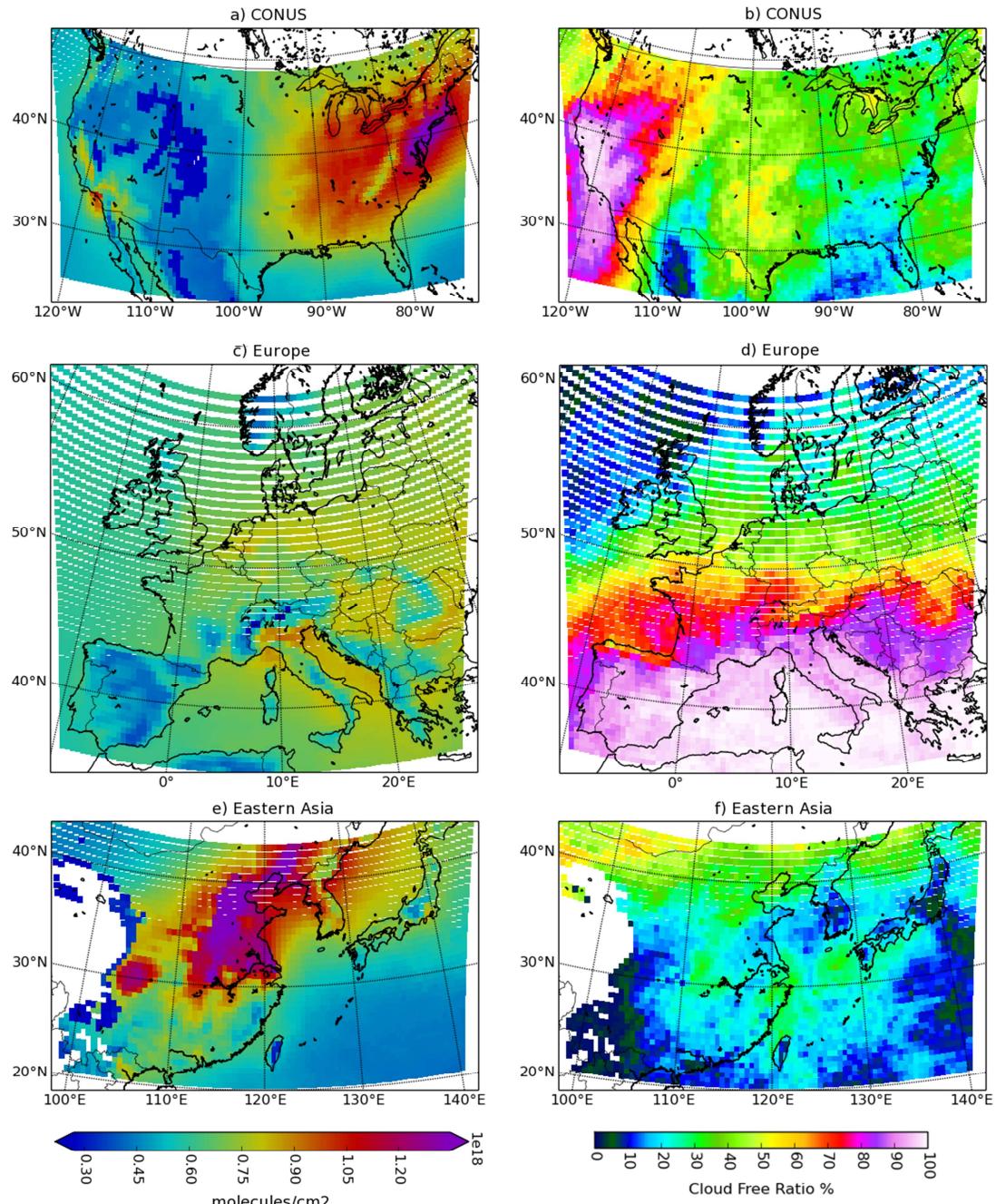
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949 **Figure 9. Snapshots of the Nature Run surface to 700 hPa partial column (a, d,**  
950 **g). Corresponding retrieved partial column (b, e, h) and corresponding**  
951 **degrees of freedom for signal (DFS) for surface to 700 hPa (c, f, j). Snapshots**  
952 **are captured at daytime but different dates following regions: 4 July 2006**  
953 **02UT CONUS, 14 July 2006 10UT Europe, 22 July 2006 18UT Eastern Asia.**  
954 **Deep colors are the cloud-free pixels. Faded colors represent cloud-**  
955 **contaminated pixels that are not used in further processing.**  
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 958 **Figure 10.** Scatter plots showing variation of degrees of freedom for signal of  
 959 surface to 700 hPa versus predictors with highest impacts to the multi-linear  
 960 regression fit (see table 2 and text for details). Red are day-time values (3pm  
 961 local time) and blue are night-time values (3am local time) 5 July 2006.  
 962 Vertical solid lines indicate the mean value of the distribution used to build  
 963 the training sets and dashed lines indicate associated  $\pm\sigma$  (standard deviation).  
 964 Dotted lines indicate associated  $+2\sigma$  for CO training set.



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 966 **Figure 11.** Left panels: scatter plots of Nature run surface-700 hPa partial  
 967 columns ( $X_t$ ) versus corresponding retrieved partial columns ( $X_r$ ). Right  
 968 panels: Smoothing error ( $X_s$ ) versus corresponding retrieved error ( $X_e$ ). Dates  
 969 are the same as described in figure 10.



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971 **Figure 12. Low-resolution observation simulations used for the assimilation**  
972 **runs. Left panels: July 2006 average retrieved CO surface-700 hPa partial**  
973 **column. Right panels: Cloud free ratio for July 2006.**

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**Highlights**

- A constellation of geostationary platforms for mapping pollutant sources and variability is described
- Observation simulation without radiative transfer model is computationally cheap
- Impacts of clouds are diagnosed and is dependent of the weather regime
- A detailed analysis of the simulated observation sensitivity is performed
- Limitations of the method are discussed